

### 1. Linearization of Ricci and scalar curvature.

We consider variations of a background metric  $b$ . All covariant derivatives, inner products and traces are with respect to  $b$  (Levi-Civita connection), unless noted otherwise. *Convention: the Laplacian of a function is the trace of its Hessian.*

*Step 1.* Let  $g_t = b + th$ ,  $h \in \text{Sym}_M^2$ . The difference of L-C connections is a (2,1) symmetric tensor:

$$\Gamma^t(X, Y) = D_X^t Y - D_X^b Y,$$

and we denote by  $\dot{\Gamma}$  its first variation. To compute the variation in curvature, define also the (3,1) tensor:

$$\Gamma^{2t}(X, Y, Z) = \Gamma(X, \Gamma(Y, Z)),$$

and then we have:

$$R^t(X, Y)Z - R^b(X, Y)Z = (D_X \Gamma^t)(Y, Z) - (D_Y \Gamma^t)(X, Z) - (\Gamma^{2t}(X, Y, Z) - \Gamma^{2t}(Y, X, Z)).$$

Since  $\Gamma = 0$  at  $t = 0$ , we have for the variation of the (3,1) Riemann tensor:

$$\dot{R}(X, Y)Z = (D_X \dot{\Gamma})(Y, Z) - (D_Y \dot{\Gamma})(X, Z).$$

The Koszul formula gives the first variation of  $\Gamma$ :

$$\langle \dot{\Gamma}(X, Y), Z \rangle = \frac{1}{2} [(D_X h)(Y, Z) + (D_Y h)(X, Z) - (D_Z h)(X, Y)].$$

*Step 2: variation of Ricci.* From now on, assume all vector fields in sight have zero covariant derivative at a fixed point  $p \in M$ . For the (4,0) Riemann tensor, we find:

$$\begin{aligned} \langle \dot{R}(X, Y)Z, W \rangle &= X(\langle \dot{\Gamma}(Y, Z), W \rangle) - Y(\langle \dot{\Gamma}(X, Z), W \rangle) \\ &= \frac{1}{2} X[(D_Y h)(Z, W) + (D_Z h)(Y, W) - (D_W h)(Y, Z)] - (Y \leftrightarrow X) \\ &= \frac{1}{2} [(D_X D_Y h)(Z, W) + (D_X D_Z h)(Y, W) - (D_X D_W h)(Y, Z)] - (Y \leftrightarrow X). \end{aligned}$$

Now take trace over  $Y, Z$ : let  $(e_i)$  be a local orthonormal frame, normal at  $p$  (sum over repeated indices implicit in what follows):

$$\begin{aligned} \langle \dot{R}(X, e_i)e_i, W \rangle &= \frac{1}{2} [(D_X D_{e_i} h)(e_i, W) + (D_x D_{e_i} h)(e_i, W) - (D_X D_W h)(e_i, e_i)] \\ &\quad - \frac{1}{2} [(D_{e_i} D_X h)(e_i, W) + (D_{e_i} D_{e_i} h)(X, W) - (D_{e_i} D_W h)(X, e_i)]. \end{aligned}$$

Now replace the last term in the second line by:  $-(D_W D_{e_i} h)(X, e_i) - (R(W, e_i)h)(X, e_i)$ . Rearranging terms, we find:

$$\begin{aligned} \dot{Ric}(X, W) &= -\frac{1}{2}(D_{e_i} D_{e_i} h)(X, W) - \frac{1}{2}(D_X D_W h)(e_i, e_i) \\ &\quad + \frac{1}{2}(D_X D_{e_i} h)(e_i, W) + \frac{1}{2}(D_W D_{e_i} h)(e_i, X) \\ &\quad + \frac{1}{2}[R(X, e_i)h](e_i, W) + \frac{1}{2}[R(W, e_i)h](e_i, X). \end{aligned}$$

*Step 3.* We now express the various terms as geometric differential operators. First we have:

$$\delta : Sym_M^2 \rightarrow \Omega^1(M), \quad (\delta h)(W) = -(D_{e_i} h)(e_i, W)$$

and its formal adjoint  $\delta^* : \Omega_M^1 \rightarrow Sym_M^2$ , the symmetrized covariant derivative:

$$(\delta^* \omega)(X, Y) = \frac{1}{2}[(D_X \omega)(Y) + (D_Y \omega)(X)], \quad \omega \in \Omega_M^1.$$

We easily find then:

$$\delta^*(\delta h)(X, W) = -\frac{1}{2}[(D_X D_{e_i} h)(e_i, W) + (D_W D_{e_i} h)(e_i, X)].$$

To understand the curvature terms, recall how  $R(X, Y)$  acts on symmetric bilinear forms:

$$\begin{aligned} [R(X, e_i)h](e_i, W) &= h(R(X, e_i)e_i, W) + h(e_i, R(X, e_i)W) \\ &= Ric(X, e_j)h(e_j, W) - h(R(e_i, X)W, e_i); \end{aligned}$$

and similarly for the term obtained by symmetrizing this in  $(X, W)$ .

Recall the symmetric product of two symmetric bilinear forms defined by:

$$(k \circ h)(X, W) = k(X, e_i)h(e_i, W), \quad k, h \in Sym_M^2$$

And also the ‘action of the Riemann tensor on symmetric bilinear forms’:

$$\mathcal{R}[h](X, W) = h(R(e_i, X)W, e_i) = h(R(e_i, W)X, e_i)$$

(cp. [Besse, 1.131(b)]; it’s like taking a Ricci trace, but using  $h$  instead of the metric.) With these definitions, the curvature terms become:

$$\frac{1}{2}[R(X, e_i)h](e_i, W) + \frac{1}{2}[R(W, e_i)h](e_i, X) = \frac{1}{2}(Ric \circ h + h \circ Ric)(X, W) - \mathcal{R}[h](X, W).$$

Putting everything together, we find for the variation of Ricci:

$$\dot{Ric}[h] = -\frac{1}{2}D_{e_i, e_i}^2 h - \frac{1}{2}Hess(tr_b h) - \delta^*(\delta h) + \frac{1}{2}[Ric \circ h + h \circ Ric] - \mathcal{R}[h].$$

(cp. [Besse, 1.180a]).

*Variation of Scalar curvature.* From  $Scal^{g_t} = tr_{g_t} Ric^{g_t}$  follows:

$$Scal[h] = -\langle h, Ric_b \rangle_b + tr_b(Ric[h]).$$

And then it turns out (easily checked) that the curvature terms in the variation of  $Ric$ , combined, have zero trace! We also have:

$$\delta^*(\delta h)(e_j, e_j) = [D_{e_j}(\delta h)](e_j) = -\delta(\delta h).$$

The traces of  $D_{e_i, e_i}^2 h$  and of  $Hess(tr_b h)$  are both equal to  $\Delta(tr_b h)$ . We conclude:

$$Scal[h] = -\Delta(tr_b h) + \delta(\delta h) - \langle Ric, h \rangle_b.$$

(cp. [Besse, 1.174e], where their convention for the Laplacian on functions has the opposite sign to ours.)

**2. Adjoints.** We have the linearization of scalar curvature at a background metric  $b$ :

$$L_b : Sym_M^2 \rightarrow C_M^\infty, \quad L_b[h] = -\Delta(tr_b h) + \delta(\delta h) - \langle Ric, h \rangle_b$$

and wish to compute its formal adjoint  $L_b^* : C_M^\infty \rightarrow Sym_M^2$ . For two of the terms, this is clear. On the other hand, if either  $V$  or  $h$  have compact support in  $M$ :

$$\int_M V \delta(\delta h) d\mu_b = \int_M \langle dV, \delta h \rangle d\mu_b = \int_M \langle \delta^*(dV), h \rangle d\mu_b = \int_M \langle Hess(V), h \rangle d\mu_b.$$

Here we used the fact  $\delta$  (on one-forms) is the formal adjoint of the exterior differential  $d$ ,  $\delta^*$  (the symmetrized covariant derivative, taking  $\Omega_M^1$  to  $Sym_M^2$ ) the formal adjoint of  $\delta : Sym_M^2 \rightarrow \Omega_M^1$ , and:

$$(\delta^* dV)(X, Y) = \frac{1}{2}[(D_X dV)Y + (D_Y dV)X] = Hess(V)(X, Y).$$

We conclude the formal adjoint is:

$$L_b^* : C_M^\infty \rightarrow Sym_M^2, \quad L_b^*[V] = Hess_b(V) - (\Delta V)b - V Ric_b.$$

Consider now the case in which neither  $V$  nor  $h$  have compact support. Pointwise, we have:

$$\langle L_b^*[V], h \rangle - \langle V, L_b[h] \rangle = \langle Hess(V), h \rangle - (\Delta V)tr_b h - V\delta(\delta h) + V\Delta(tr_b h).$$

$$-\delta(\delta h) + \Delta(tr_b h) = \delta\mu, \quad \mu = -\delta h - d(tr_b h) \in \Omega_M^1.$$

In particular, if  $b$  is Ricci-flat, the constants (say  $V \equiv 1$ ) are in  $Ker L_b^*$ , and we have, integrating over the compact manifold with boundary  $M$ :

$$\int_M L_b[h] d\mu_b = - \int_M (\delta\mu) d\mu_b = \int_M div_b(\mu^\#) = \int_{\partial M} \mu[\nu_b] d\sigma_b,$$

by the divergence theorem ( $\nu_b$  is the unit outward normal in the background metric). This relates the bulk integral of linearized scalar curvature to the boundary integral of the ADM mass integrand.  $\mu[\nu_b]$

In the general case, we have:

$$\delta[V(-\delta h - d(tr_b h)) - i_{\nabla V} h + (tr_b h)dV] = \\ \langle dV, \delta h \rangle - V\delta(\delta h) + \langle dV, d(tr_b h) \rangle - V\delta d(tr_b h) - \delta(i_{\nabla V} h) - \langle d(tr_b h), dV \rangle - (tr_b h)\Delta V.$$

Now use:

$$\delta(i_{\nabla V} h) = (\delta h)(\nabla V) - \langle Hess(V), h \rangle.$$

After cancelation, we find the pointwise relation:

$$\delta(\mu_{(V,h)}) = \langle L_b^*[V], h \rangle - \langle V, L_b[h] \rangle, \quad \mu_{(V,h)} = V(-\delta h - d(tr_b h)) - i_{\nabla V} h + (tr_b h)dV \in \Omega_M^1.$$

Integrating over the compact manifold with boundary  $M$

$$\int_M (\langle V, L_b[h] \rangle - \langle L_b^*[V], h \rangle) d\mu_b = \int_M div_b(\mu_{(V,h)}^\#) d\mu_b = \int_{\partial M} \mu_{(V,h)}[\nu_b] d\sigma_b.$$

### 3. Variation of the Einstein-Hilbert functional.

We consider the first variation of:

$$\mathcal{R}_b = \int_M S_b d\mu_b$$

under the variation of metric:  $g^t = b + th$ ,  $h \in Sym_M^2$ . We have:

$$\dot{S} = -\Delta(tr_b h) + \delta(\delta h) - \langle Ric, h \rangle_b, \quad (\dot{d}\mu) = \frac{1}{2}(tr_b h)d\mu_b.$$

Thus:

$$\dot{\mathcal{R}} = \int_M [\delta(\delta h + d(tr_b h)) - \langle G^b, h \rangle] d\mu_b, \quad G^b = Ric^b - \frac{S_b}{2}b,$$

the ‘Einstein tensor’ of  $b$ . Using the divergence theorem:

$$\dot{\mathcal{R}} = - \int_M \langle G^b, h \rangle d\mu_b + \int_{\partial M} \mu_h[\nu_b] d\sigma_b, \quad \mu_h = -\delta h - d(tr_b h) \in \Omega_M^1.$$

We conclude the critical metrics for  $\mathcal{R}$  (under variations  $h$  with compact support) are those with vanishing Einstein tensor  $G$ . And for metrics with vanishing  $G$  and variation tensor  $h = \dot{g}$ , we have the suggestive relation:

$$\dot{\mathcal{R}} - \dot{m} = 0, \quad m = \int_{\partial M} \mu_g[\nu_b] d\sigma_b.$$

**4. Where the mass comes from.** (cp. [Michel] and [Herzlich].)

We consider  $(M, g)$  complete noncompact (with one end, for simplicity), asymptotic to a ‘background’  $(\mathbb{M}_0, b)$  (typically euclidean or hyperbolic  $n$ -space, in the following sense: there exists a chart from an exterior region (complement of a ball)  $E_0 \subset \mathbb{M}_0$  to  $M \setminus K$ ,  $K \subset M$  compact:  $\phi : E_0 \rightarrow M \setminus K$ , and letting  $g_\phi = \phi^*g$ , we have on  $E_0$ :

$$h = g_\phi - b = O_2(r^{-\tau}),$$

where  $r$  is distance in  $\mathbb{M}_0$  to a fixed point. Consider the space of ‘static potentials’ on  $b$ :

$$\mathcal{N}_b = \{V \in C^\infty(\mathbb{M}_0; L_b^*[V] = 0\}, \quad L_b^*(V) = Hess(V) - (\Delta_b V)b - VRic_b,$$

the formal adjoint of the linearization at  $b$  of the scalar curvature map. Consider the Taylor theorem representation:

$$Scal(g_\phi) - Scal(b) = L_b[h] + Q_b(h),$$

where  $Q_b(h)$  is the quadratic (and higher order) remainder.

*Exercise.* If  $V \in \mathcal{N}_b$ , the metric  $V^2 dt \oplus b$  on  $\mathbb{R} \times \mathbb{M}_0$  is Ricci-flat.

The main assumption is  $g_\phi$  is asymptotic to  $b$  as  $r \rightarrow \infty$ , at a fast enough rate that:

- (1)  $\langle V, Scal(g_\phi) - Scal(b) \rangle_b$  is  $d\mu_b$ -integrable;
- (2)  $Q_b(V, h) = \langle V, Q_b(h) \rangle_b$  is  $d\mu_b$ -integrable, for all  $V \in \mathcal{N}_b$ .

If  $V \in \mathcal{N}_b$ , pointwise on  $E_0$  we have:

$$\begin{aligned} \langle V, Scal(g_\phi) - Scal(b) \rangle &= \langle V, L_b[h] \rangle + \langle V, Q_b(h) \rangle \\ &= -\delta_b(\mu^\phi(V, h)) + \langle V, Q_b(h) \rangle_b, \end{aligned}$$

for a one-form  $\mu^\phi(V, h)$  on  $E_0$  given in the previous section (linear in  $V, h$  and their first order derivatives.)

Let  $(B_k)$  be an increasing exhaustion of  $\mathbb{M}_0$  by Smoothly bounded domains,  $\partial B_k = S_k$  (so for  $k \geq k_1$  large enough,  $S_k \subset E_0$ ). The divergence theorem implies:

$$\oint_{S_k} \mu^\phi(V, h)[\nu_b] d\sigma_b = \oint_{S_{k_1}} \mu^\phi(V, h)[\nu_n] d\sigma_b + \int_{B_k \setminus B_{k_1}} [\langle V, Scal(g_\phi - Scal(b)) - Q_b(V, h) \rangle] d\mu_b.$$

The integrability conditions (1) and (2) guarantee the RHS has a limit as  $k \rightarrow \infty$ , and that this limit is independent of  $k_1$ . We conclude the limit:

$$\lim_k \oint_{S_k} \mu^\phi(V, h)[\nu_b] d\sigma_b$$

exists, and is independent of the exhaustion considered.

*Remark:* Additional conditions are needed to ensure the limit is independent of the chart  $\phi$ . Roughly speaking, they are: (i)  $Scal(b)$  is constant and (ii) Any two charts  $\phi, \psi$  with the property that  $g_\phi, g_\psi$  are asymptotic to  $b$  differ by a diffeomorphism (of some exterior region  $E_R$ ) whose leading term is an isometry of  $b$ . (This is a ‘rigidity condition’ on  $b$ ; for details, see [Michel].)

*Example 1: asymptotically flat manifolds.*  
 $(\mathbb{M}_0, b) = (\mathbb{R}^n, \delta)$ : euclidean space.

We assume  $h = g - \delta = O_2(r^{-\tau})$  on  $E_0$ . The remainder term satisfies:

$$|Q(1, e)| \leq C(|\partial h|^2 + |h|_b |\partial^2 h|) = O(r^{-2\tau-2}).$$

Thus the integrability conditions (1), (2) above are satisfied if  $Scal_g \in L^1(M, g)$  and  $2\tau + 2 > n$ , or  $\tau > \frac{n-2}{2}$ .

Since  $L_0^*[V] = Hess_0(V) - (\Delta_0 V)\delta$  (euclidean Hessian and Laplacian), taking traces one sees easily that  $V \in \mathcal{N}_b$  iff  $Hess_0(V) = 0$ , or  $V$  is affine (linear plus constant):  $\mathcal{N}_b$  is  $(n+1)$ -dimensional. For  $V = 1$ , the mass one-form is:

$$\mu(1, h) = \sum_{i,j} (\partial_i h_{ij} - \partial_j h_{ii}) dx^j \in \Omega^1(E_0),$$

and the mass is:

$$m(g) = \lim_k \oint \mu(1, h)[\nu_0] d\sigma_0,$$

up to a normalization constant depending only on  $n$ .

*Example 2: asymptotically hyperbolic manifolds.*  $(\mathbb{M}_0, b) = (\mathbb{H}^n, g_{hyp})$ , hyperbolic space.:  $g_{hyp} = dr^2 + \sinh^2 r g_{S^{n-1}}$ , the standard metric on the sphere. The asymptotic conditions are:

$$|h| = O_2(e^{-r\tau}), \quad h = \phi^* g - b.$$

The static potentials  $V$  (kernel of  $L_b^*$ ) are solutions of:

$$Hess_b(V) - (\Delta_b V)b = V(Ric_b) = -(n-1)Vb.$$

It is easy to see (*exercise*) this is equivalent to  $Hess_b(V) = Vb$ . As in the euclidean case, the space  $\mathcal{N}_b$  of solutions has dimension  $n+1$ :

$$\mathcal{N}_b = span\{V_0, V_1, \dots, V_n\}, \quad V_0 = \cosh r, V_i = (\sinh r)\omega_i,$$

where  $\omega = (\omega_1, \dots, \omega_n)$  is the standard embedding of  $S^{n-1}$  into  $R^n$ .

Consider the integrability conditions (1) and (2): since  $|V| + |\nabla V| = O(e^r)$ , we need  $e^r(Scal(\phi^* g) + n(n-1))$  integrable; while  $|Q_b(V, h)| = O(e^r e^{-2\tau r})$ ; since  $d\mu_b = O(e^{(n-1)r})$ , we have  $Q_b(V, h)d\mu_b = O(e^{nr} e^{-2\tau r})$ , so for integrability we need:  $\tau > n/2$ . Under these conditions, the limit:

$$m_g(V) = \lim_k \oint_{S_k} \mu_b(V, h)[\nu_b] d\sigma_n, \quad \mu_b(V, h) = V(-\delta_b h - d(tr_b h)) - i_{\nabla_b V} h + (tr_b h)dV \in \Omega^1(E_0)$$

exists, and defines a linear functional on  $\mathcal{N}_b$ .